

Master programmes in Artificial Intelligence 4 Careers in Europe



#### University of Cyprus – MSc Artificial Intelligence

# **MAI644 – COMPUTER VISION Lecture 16: Stereo Vision**

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#### Last time

Connecting Europe Facility

- Full camera model in matrix form
- Camera calibration
- Calibration Projective camera model
- Calibration Affine camera model









# Today's Agenda

- Recovery of world position
- Triangulation
- Epipolar Geometry











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- Previously we saw that the imaging process can be described as a transformation in homogeneous coordinates.
- If we can invert this transformation, the world coordinates of each pixel in the image can be computed.
- Recovering world coordinates of objects based on the projection on an image is known as shape recovery or depth recovery.









- Is this possible using a single camera?
- The camera needs to be calibrated, i.e. we know all its parameters
- Remember the projective camera model:

$$
\widetilde{w} = P\widetilde{X}
$$
  
\n
$$
\Leftrightarrow \begin{bmatrix} su \\ sv \\ s \end{bmatrix} = \begin{bmatrix} p_{11} & p_{12} & p_{13} & p_{14} \\ p_{21} & p_{22} & p_{23} & p_{24} \\ p_{31} & p_{32} & p_{33} & p_{34} \end{bmatrix} \begin{bmatrix} X \\ Y \\ Z \\ Z \\ 1 \end{bmatrix}
$$

• Unfortunately the transformation described in  $P$  is not invertible and the world point  $X$  cannot be uniquely determined





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# Recovery of world position

#### • Depth ambiguity



Courtesy slide S. Lazebnik









• Each observed feature on the image gives 2 equations with 3 unknowns and therefore defines a line (a ray) of solutions for  $X$ 









- Each observed feature on the image gives 2 equations with 3 unknowns and therefore defines a line (a ray) of solutions for  $X$
- This system of equations is under-constrained.
- This can be seen by the size of *. There are more columns than rows.*

$$
\widetilde{w} = P\widetilde{X}
$$
  
\n
$$
\Leftrightarrow \begin{bmatrix} su \\ sv \\ s \end{bmatrix} = \begin{bmatrix} p_{11} & p_{12} & p_{13} & p_{14} \\ p_{21} & p_{22} & p_{23} & p_{24} \\ p_{31} & p_{32} & p_{33} & p_{34} \end{bmatrix} \begin{bmatrix} X \\ Y \\ Z \\ Z \\ 1 \end{bmatrix}
$$









- Under-constrained problems never have a unique solution.
- To uniquely recover  $X$ , additional views must be used, so that the transformation between  $w$  (pixel coordinates) and  $X$  (world coordinates) is *forced* to become invertible.
- This is the subject of stereo vision.









• Two eyes/cameras help.











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- In stereo vision at least two cameras are set up to view the  $3D$  scene.
- Each 3D world location X projects to pixel w on camera 1 (O) and to pixel  $w'$  on camera 2  $(O')$ .









• If both cameras are calibrated, the  $3D$  world location  $\boldsymbol{X}$  projected on the pair of corresponding pixel locations  $w$  and  $w'$  can be estimated via a process known as triangulation.











• Consider the projection of  $X$  onto  $w$ :

$$
\widetilde{w} = P\widetilde{X}
$$
  
\n
$$
\Leftrightarrow \begin{bmatrix} su \\ sv \\ s \end{bmatrix} = \begin{bmatrix} p_{11} & p_{12} & p_{13} & p_{14} \\ p_{21} & p_{22} & p_{23} & p_{24} \\ p_{31} & p_{32} & p_{33} & p_{34} \end{bmatrix} \begin{bmatrix} X \\ Y \\ Z \\ Z \\ 1 \end{bmatrix}
$$

• Compute the equation for  $u$  and rearrange so our unknowns  $X, Y, Z$ are on the left:

$$
u = \frac{su}{s} = \frac{p_{11}X + p_{12}Y + p_{13}Z + p_{14}}{p_{31}X + p_{32}Y + p_{33}Z + p_{34}}
$$
  
\n
$$
\Rightarrow p_{11}X + p_{12}Y + p_{13}Z + p_{14} = p_{31}uX + p_{32}uY + p_{33}uZ + p_{34}u
$$
  
\n
$$
\Rightarrow (p_{11} - p_{31}u)X + (p_{12} - p_{32}u)Y + (p_{13} - p_{33}u)Z = p_{34}u - p_{14}
$$









• Compute the equation for  $u$  and rearrange so our unknowns  $X, Y, Z$ are on the left:

$$
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$$
  
\n
$$
\Rightarrow p_{11}X + p_{12}Y + p_{13}Z + p_{14} = p_{31}uX + p_{32}uY + p_{33}uZ + p_{34}u
$$
  
\n
$$
\Rightarrow (p_{11} - p_{31}u)X + (p_{12} - p_{32}u)Y + (p_{13} - p_{33}u)Z = p_{34}u - p_{14}
$$

• Put this in matrix form:

$$
[p_{11} - p_{31}u \quad p_{12} - p_{32}u \quad p_{13} - p_{33}u] \begin{bmatrix} X \\ Y \\ Z \end{bmatrix} = [p_{34}u - p_{14}]
$$
  

$$
\Leftrightarrow Ax = b
$$











• Put this in matrix form:

$$
[p_{11} - p_{31}u \quad p_{12} - p_{32}u \quad p_{13} - p_{33}u] \begin{bmatrix} X \\ Y \\ Z \end{bmatrix} = [p_{34}u - p_{14}]
$$
  

$$
\Leftrightarrow Ax = b
$$

- $\bullet$  *A* is a 1 $x3$  matrix
- The resulting system is under-constrained











• Put this in matrix form:

$$
[p_{11} - p_{31}u \quad p_{12} - p_{32}u \quad p_{13} - p_{33}u] \begin{bmatrix} X \\ Y \\ Z \end{bmatrix} = [p_{34}u - p_{14}]
$$
  

$$
[p_{21} - p_{31}v \quad p_{22} - p_{32}v \quad p_{23} - p_{33}v] \begin{bmatrix} X \\ Y \\ Z \end{bmatrix} = [p_{34}v - p_{24}]
$$
  

$$
\Leftrightarrow Ax = b
$$

• By computing  $v$  in the same way as  $u$  and rearranging we can add a new row in  $\boldsymbol{A}$  and in  $\boldsymbol{b}$ , making it 2x3









• Put this in matrix form:

$$
[p'_{11} - p'_{31}u' \quad p'_{12} - p'_{32}u' \quad p'_{13} - p'_{33}u'] \begin{bmatrix} X \\ Y \\ Z \end{bmatrix} = [p'_{34}u' - p'_{14}]
$$
  

$$
[p'_{21} - p'_{31}v' \quad p'_{22} - p'_{32}v' \quad p'_{23} - p'_{33}v'] \begin{bmatrix} X \\ Y \\ Z \end{bmatrix} = [p'_{34}v' - p'_{24}]
$$
  

$$
\Leftrightarrow Ax = b
$$

• By also considering the projection of  $X$  onto  $w'$ , a new pair of rows will be added to  $A$ , thus forcing it to be  $4x3$ , i.e. over-constrained









• Here is the resulting system

$$
\begin{bmatrix} p_{11} - p_{31}u & p_{12} - p_{32}u & p_{13} - p_{33}u \\ p_{21} - p_{31}v & p_{22} - p_{32}v & p_{23} - p_{33}v \\ p'_{11} - p'_{31}u' & p'_{12} - p'_{32}u' & p'_{13} - p'_{33}u' \\ p'_{21} - p'_{31}v' & p'_{22} - p'_{32}v' & p'_{23} - p'_{33}v' \end{bmatrix} \begin{bmatrix} X \\ Y \\ Z \end{bmatrix} = \begin{bmatrix} p_{34}u - p_{14} \\ p_{34}v - p_{24} \\ p'_{34}u' - p'_{14} \\ p'_{34}v' - p'_{24} \end{bmatrix}
$$

• Where  $p'_{11}$  etc. are the parameters inside the camera projection matrix for camera 2 ( $O'$ ), and  $w' = (u', v')$  are the pixel coordinates of  $X$  projected on the image plane  $I'$  of the second camera









 $\bullet$  Using an additional camera has forced  $\bm{A}$  to become over-constrained.

$$
\begin{bmatrix} p_{11} - p_{31}u & p_{12} - p_{32}u & p_{13} - p_{33}u \\ p_{21} - p_{31}v & p_{22} - p_{32}v & p_{23} - p_{33}v \\ p'_{11} - p'_{31}u' & p'_{12} - p'_{32}u' & p'_{13} - p'_{33}u' \\ p'_{21} - p'_{31}v' & p'_{22} - p'_{32}v' & p'_{23} - p'_{33}v' \end{bmatrix} \begin{bmatrix} X \\ Y \\ Z \end{bmatrix} = \begin{bmatrix} p_{34}u - p_{14} \\ p_{34}v - p_{24} \\ p'_{34}u' - p'_{14} \\ p'_{34}v' - p'_{24} \end{bmatrix}
$$

• Therefore we can now find a least-squares solution:

$$
Ax = b
$$
  

$$
\Leftrightarrow x = (A^T A)^{-1} A^T b
$$









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#### Beyond triangulation

We have seen the simplest form of stereo vision: Given a pair of *calibrated* cameras observing a single feature at *corresponding* pixel locations  $w \leftrightarrow w'$ , the 3D position of the corresponding world location  $\boldsymbol{X}$  can be estimated via triangulation











#### Beyond triangulation

How is the correspondence problem solved if there are several points  $W_i\}_{i=1}^{N_1}$  $N_{1}$ in image 1 and several points  $\{ {\boldsymbol{w}'}_{\boldsymbol{j}} \}$  $j=1$  $N_2$ in image 2 ?











#### Beyond triangulation

SIFT will give us a set of proposed correspondences  $\{w_i \leftrightarrow w'_j\}$  but there will be many outliers in these proposals

The question is then how can we remove outliers ?

Using the epipolar constraint



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#### Epipolar Geometry

- To understand the epipolar constraint, we first need to understand the geometry that relates the
	- cameras
	- points in  $3D$  space
	- and their corresponding observations  $\{ {\boldsymbol{w}}_{{\boldsymbol{i}}} \leftrightarrow {\boldsymbol{w}}'{}_{{\boldsymbol{j}}}$



• This type of geometry is referred to as the epipolar geometry







Lets revisit our stereo pair











The **baseline**  $l<sub>b</sub>$  is the line joining the two optical centers.











The epipolar plane  $\Pi$  is the plane defined by the 3D point X and the optical centers of the cameras.











An epipole is the point of intersection of the baseline with the image plane. There are two epipoles  $e$  and  $e'$ , one for each image.











An epipolar line is a line of intersection of the epipolar plane with an image plane. It is the image, in one camera, of the ray from the other camera's optical centre to the point  $X$ .











For different world points  $X$ , the epipolar plane rotates about the baseline. All epipolar lines intersect at their corresponding epipole.











The **epipolar constraint** limits the search for correspondences, from the region of the whole image, to only the pixels spanned by the epipolar line.











If a point feature  $w$  is observed in one image, then its location  $w'$  in the other image must lie on its corresponding epipolar line  $l'$ 











So a simple algorithm for determining correspondences is to match each feature  $w_i$  from camera 1 to a feature  $w^\prime{}_j$  from camera 2 which is close to the epipolar line of camera 2, provided that the SIFT descriptors for  $w_i$  and  $w^\prime{}_j$  are similar. epiline









Therefore, we must calculate the equation of the epipolar lines.





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• Firstly, lets assume that the 1<sup>st</sup> camera is located at the world origin  $\boldsymbol{O}$ 

• This means that every point  $X$  is expressed in the camera coordinates of the 1<sup>st</sup> camera, since its optical center coincides with the world origin  $\Rightarrow$   $X \rightarrow X_c$ 









• If this is the case, the 2<sup>nd</sup> camera is at a location  $\boldsymbol{O}'$  in world coordinates, which can be expressed by a translation  $t$  and rotation  $R$ , w.r.t. 1<sup>st</sup> camera, i.e., the world origin











- Every 3D point  $X$  can be expressed as  $X_c$ , which is the cameracentered coordinate system of the 1<sup>st</sup> camera, and as  $X_c'$ , which is the camera-centered coordinate system of the 2<sup>nd</sup> camera.
- Since  $X = X_c$ , we can relate  $X \leftrightarrow X_c'$  or  $X_c \leftrightarrow X_c'$  using a Euclidean transformation composed by the translation  $\boldsymbol{t}$  and rotation  $\boldsymbol{R}$  of the 2<sup>nd</sup> camera, w.r.t. the 1<sup>st</sup> camera









Here is how to find an expression for the epipolar line:

$$
\widetilde{X'_c} = P_e \widetilde{X_c}
$$
\n
$$
\Leftrightarrow X'_c = R X_c + t
$$
\n
$$
\Leftrightarrow t \times X'_c = t \times R X_c + t \times t^0
$$
\n
$$
\Leftrightarrow t \times X'_c = t \times R X_c + t \times t^0
$$
\n
$$
\Leftrightarrow X'_{c} \cdot (t \times X'_{c}) = X'_{c} \cdot (t \times R X_{c})
$$
\n
$$
\Leftrightarrow 0 = X'_{c} \cdot (t \times R X_{c})
$$
\n
$$
\Leftrightarrow 0 = X'_{c} \cdot (t \times R X_{c})
$$
\n
$$
\Leftrightarrow 0 = X'_{c} \cdot (t \times R X_{c})
$$







This can be rewritten in matrix form:

 $X'_{c} \cdot (t \times RX_{c}) = 0$  $\Leftrightarrow X_c^{\prime T} E X_c = 0$ 

Where  $\mathbf{E} = \mathbf{T}_{\times} \mathbf{R}$  is the essential matrix, and

$$
\boldsymbol{T}_{\times} = \begin{bmatrix} 0 & -t_z & t_y \\ t_z & 0 & -t_x \\ -t_y & t_x & 0 \end{bmatrix}
$$

is a matrix representing the cross product with **t** such that  $\mathbf{t} \times \mathbf{v} = T_{\times} \mathbf{v}$ 







#### Fundamental Matrix

Recall that for  $X_c \leftrightarrow w$ :  $\widetilde{w} = K X_c \Leftrightarrow X_c = K^{-1} \widetilde{w}$ 

and similarly, for  $X'_c \leftrightarrow w' \colon \hat{w}' = K'X'_c \Leftrightarrow X'_c = K'^{-1}\tilde{w}'$ 

Combining the two equations yields the equation of the two epipolar lines in pixel coordinates:

$$
X_c^{\prime T} \mathbf{E} \mathbf{X_c} = 0
$$
  
\n
$$
\Rightarrow (\mathbf{K'}^{-1} \widetilde{\mathbf{w}}')^T \mathbf{E} (\mathbf{K}^{-1} \widetilde{\mathbf{w}}) = 0
$$
  
\n
$$
\Rightarrow \widetilde{\mathbf{w}}^{\prime T} (\mathbf{K'}^{-T} \mathbf{E} \mathbf{K}^{-1}) \widetilde{\mathbf{w}} = 0
$$
  
\n
$$
\Rightarrow \widetilde{\mathbf{w}}^{\prime T} \mathbf{F} \widetilde{\mathbf{w}} = 0
$$

where  $\boldsymbol{F} = \boldsymbol{K}'^{-T}\boldsymbol{E}\boldsymbol{K}^{-1}$  is the fundamental matrix.









#### Fundamental Matrix

This is the equation of the epipolar line in either camera:

 $\widetilde{w}^{\prime T} F \ \widetilde{w} = 0$ 

Assuming we know the fundamental matrix, for every point  $w$  in image 1, this expression gives us the line in image 2 on which the corresponding ′ must lie, and *vice versa.*

- $l' = F \widetilde{w}$  is the epipolar line in the 2<sup>nd</sup> image, associated with  $\bm{w}$
- $l = \boldsymbol{F}^T \boldsymbol{\widetilde{w}}'$  is the epipolar line in the 1<sup>st</sup> image, associated with  $\boldsymbol{w}'$
- $l_i: ax + by + c = 0$









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#### Examples

#### Here are a few examples of the epipolar constraint













Epipolar constraint examples: Parallel image planes



- Baseline intersects the image planes at infinity
- Epipoles are at infinity

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• Epipolar lines are parallel to the  $u$ -axis of each image plane









#### Epipolar constraint examples: Parallel image planes











Epipolar constraint examples: Forward translation



• The **epipoles** have the same position in both images









#### Epipolar constraint examples: Forward translation







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#### Fundamental Matrix

In order to apply the epipolar constraint we need to know the fundamental matrix:

> $F = K'^{-T} E K^{-1}$  $= K'^{-T}T_{\times}RK^{-1}$

All the parameters of  $\bm{F}$  come from the calibration of the two cameras.

Remember that the perspective camera model  $P_{ps} = K[R|T]$  contains this information.

If, however, these are not available (e.g., because we used a projective camera model to calibrate our cameras),  $\bm{F}$  must be estimated using known image correspondences.











To estimate the fundamental matrix, we follow a similar approach as in camera calibration

$$
\widetilde{\mathbf{w}}^{T} \mathbf{F} \; \widetilde{\mathbf{w}} = 0 \Rightarrow [u' \; v' \; 1] \begin{bmatrix} f_{11} & f_{12} & f_{13} \\ f_{21} & f_{22} & f_{23} \\ f_{31} & f_{32} & f_{33} \end{bmatrix} \begin{bmatrix} u \\ v \\ 1 \end{bmatrix} = 0
$$

Each point correspondence  $w \leftrightarrow w'$  generates a single equation for estimating the parameters inside  $\bm{F}$ .







Here are *N* such correspondences:  
\n
$$
\begin{bmatrix}\nu_1 u'_1 & v_1 u'_1 & u'_1 & u_1 v'_1 & v_1 v'_1 & v'_1 & u_1 & v_1 \\
u_N u'_N & v_N u'_N & u'_N & u_N v'_N & v_N v'_N & v'_N & u_N & v_N\n\end{bmatrix}\n\begin{bmatrix}\nf_{11} \\
f_{12} \\
f_{21} \\
f_{22} \\
f_{23} \\
f_{31} \\
f_{32}\n\end{bmatrix} =\n\begin{bmatrix}\n-1 \\
\vdots \\
-1\n\end{bmatrix}
$$

The minimal number for N is 8, because  $\bm{F}$  has 8 degrees of freedom (we can set  $f_{33} = 1$ )



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• In practise, 8 clearly indicated correspondences are never available

• What is available is a set of correspondences proposed by SIFT, for a large set of features, in which there are always errors









- We can use RANSAC to solve this problem
	- 1. Obtain 8 random SIFT correspondences. Use them to estimate  $\bm{F}$ .
	- 2. For every feature  $w_i$  in image 1 calculate the epipolar line in image 2. Check if the corresponding feature  $\bm{w'}_{\bm{j}}$  in image 2 proposed by SIFT falls **"close"** to the epipolar line. Count the number  $S$  of such "inliers".
	- 3. If  $S \geq T$  where T is a threshold, then there is consensus with the random sample taken in the first step. Calculate  $\bm{F}$  for all inliers and terminate here.
	- 4. If  $S < T$  then no consensus is reached. Repeat from step 1.
	- 5. If after  $N$  iterations no consensus is reached, select the model that gave the highest S, calculate F using all inliers in S and terminate







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# **Thank you.**



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