

Master programmes in Artificial Intelligence 4 Careers in Europe



University of Cyprus – MSc Artificial Intelligence

MAI644 – COMPUTER VISION Lecture 16: Stereo Vision

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Last time

- Full camera model in matrix form
- Camera calibration
- Calibration Projective camera model
- Calibration Affine camera model









Today's Agenda

- Recovery of world position
- Triangulation
- Epipolar Geometry

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- Previously we saw that the imaging process can be described as a transformation in homogeneous coordinates.
- If we can invert this transformation, the world coordinates of each pixel in the image can be computed.
- Recovering world coordinates of objects based on the projection on an image is known as **shape recovery** or **depth recovery**.









- Is this possible using a single camera ?
- The camera needs to be calibrated, i.e. we know all its parameters
- Remember the projective camera model:

$$\widetilde{\boldsymbol{w}} = \boldsymbol{P}\widetilde{\boldsymbol{X}}$$

$$\Leftrightarrow \begin{bmatrix} su\\sv\\s \end{bmatrix} = \begin{bmatrix} p_{11} & p_{12} & p_{13} & p_{14}\\p_{21} & p_{22} & p_{23} & p_{24}\\p_{31} & p_{32} & p_{33} & p_{34} \end{bmatrix} \begin{bmatrix} \boldsymbol{X}\\\boldsymbol{Y}\\\boldsymbol{Z}\\1 \end{bmatrix}$$

• Unfortunately the transformation described in ${\it P}$ is not invertible and the world point ${\it X}$ cannot be uniquely determined







• Depth ambiguity



Courtesy slide S. Lazebnik









• Each observed feature on the image gives 2 equations with 3 unknowns and therefore defines a line (a ray) of solutions for X









- Each observed feature on the image gives 2 equations with 3 unknowns and therefore defines a line (a ray) of solutions for ${\pmb X}$
- This system of equations is under-constrained.
- This can be seen by the size of \boldsymbol{P} . There are more columns than rows.

$$\widetilde{\boldsymbol{w}} = \boldsymbol{P}\widetilde{\boldsymbol{X}}$$

$$\Leftrightarrow \begin{bmatrix} su\\sv\\s \end{bmatrix} = \begin{bmatrix} p_{11} & p_{12} & p_{13} & p_{14}\\p_{21} & p_{22} & p_{23} & p_{24}\\p_{31} & p_{32} & p_{33} & p_{34} \end{bmatrix} \begin{bmatrix} \boldsymbol{X}\\\boldsymbol{Y}\\\boldsymbol{Z}\\1 \end{bmatrix}$$









- Under-constrained problems never have a unique solution.
- To uniquely recover X, additional views must be used, so that the transformation between w (pixel coordinates) and X (world coordinates) is *forced* to become invertible.
- This is the subject of **stereo vision**.







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Recovery of world position

• Two eyes/cameras help.











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- In stereo vision at least two cameras are set up to view the 3D scene.
- Each 3D world location X projects to pixel w on camera 1 (O) and to pixel \boldsymbol{w}' on camera 2 (\boldsymbol{O}').





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• If both cameras are calibrated, the 3D world location X projected on the pair of corresponding pixel locations w and w' can be estimated via a process known as triangulation.









• Consider the projection of **X** onto **w**:

$$\widetilde{\boldsymbol{w}} = \boldsymbol{P}\widetilde{\boldsymbol{X}}$$

$$\Leftrightarrow \begin{bmatrix} su\\sv\\s \end{bmatrix} = \begin{bmatrix} p_{11} & p_{12} & p_{13} & p_{14}\\p_{21} & p_{22} & p_{23} & p_{24}\\p_{31} & p_{32} & p_{33} & p_{34} \end{bmatrix} \begin{bmatrix} \boldsymbol{X}\\\boldsymbol{Y}\\\boldsymbol{Z}\\1 \end{bmatrix}$$

• Compute the equation for *u* and rearrange so our unknowns *X*, *Y*, *Z* are on the left:

$$u = \frac{su}{s} = \frac{p_{11}X + p_{12}Y + p_{13}Z + p_{14}}{p_{31}X + p_{32}Y + p_{33}Z + p_{34}}$$

$$\Rightarrow p_{11}X + p_{12}Y + p_{13}Z + p_{14} = p_{31}uX + p_{32}uY + p_{33}uZ + p_{34}u$$

$$\Rightarrow (p_{11} - p_{31}u)X + (p_{12} - p_{32}u)Y + (p_{13} - p_{33}u)Z = p_{34}u - p_{14}$$







• Compute the equation for *u* and rearrange so our unknowns *X*, *Y*, *Z* are on the left:

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$$\Rightarrow (p_{11} - p_{31}u)X + (p_{12} - p_{32}u)Y + (p_{13} - p_{33}u)Z = p_{34}u - p_{14}$$

• Put this in matrix form:

$$\begin{bmatrix} p_{11} - p_{31}u & p_{12} - p_{32}u & p_{13} - p_{33}u \end{bmatrix} \begin{bmatrix} X \\ Y \\ Z \end{bmatrix} = \begin{bmatrix} p_{34}u - p_{14} \end{bmatrix}$$
$$\Leftrightarrow Ax = b$$







• Put this in matrix form:

$$\begin{bmatrix} p_{11} - p_{31}u & p_{12} - p_{32}u & p_{13} - p_{33}u \end{bmatrix} \begin{bmatrix} X \\ Y \\ Z \end{bmatrix} = \begin{bmatrix} p_{34}u - p_{14} \end{bmatrix}$$
$$\Leftrightarrow Ax = b$$

- \boldsymbol{A} is a 1x3 matrix
- The resulting system is under-constrained









• Put this in matrix form:

$$\begin{bmatrix} p_{11} - p_{31}u & p_{12} - p_{32}u & p_{13} - p_{33}u \end{bmatrix} \begin{bmatrix} X \\ Y \\ Z \end{bmatrix} = \begin{bmatrix} p_{34}u - p_{14} \end{bmatrix}$$
$$\begin{bmatrix} p_{21} - p_{31}v & p_{22} - p_{32}v & p_{23} - p_{33}v \end{bmatrix} \begin{bmatrix} X \\ Y \\ Z \end{bmatrix} = \begin{bmatrix} p_{34}v - p_{24} \end{bmatrix}$$
$$\Leftrightarrow \mathbf{A}\mathbf{x} = \mathbf{b}$$

• By computing v in the same way as u and rearranging we can add a new row in A and in b, making it 2x3







• Put this in matrix form:

$$\begin{bmatrix} p'_{11} - p'_{31}u' & p'_{12} - p'_{32}u' & p'_{13} - p'_{33}u' \end{bmatrix} \begin{bmatrix} X \\ Y \\ Z \end{bmatrix} = \begin{bmatrix} p'_{34}u' - p'_{14} \end{bmatrix}$$
$$\begin{bmatrix} p'_{21} - p'_{31}v' & p'_{22} - p'_{32}v' & p'_{23} - p'_{33}v' \end{bmatrix} \begin{bmatrix} X \\ Y \\ Z \end{bmatrix} = \begin{bmatrix} p'_{34}v' - p'_{24} \end{bmatrix}$$
$$\Leftrightarrow Ax = b$$

• By also considering the projection of X onto w', a new pair of rows will be added to A, thus forcing it to be 4x3, i.e. over-constrained









• Here is the resulting system

$$\begin{bmatrix} p_{11} - p_{31}u & p_{12} - p_{32}u & p_{13} - p_{33}u \\ p_{21} - p_{31}v & p_{22} - p_{32}v & p_{23} - p_{33}v \\ p'_{11} - p'_{31}u' & p'_{12} - p'_{32}u' & p'_{13} - p'_{33}u' \\ p'_{21} - p'_{31}v' & p'_{22} - p'_{32}v' & p'_{23} - p'_{33}v' \end{bmatrix} \begin{bmatrix} X \\ Y \\ Z \end{bmatrix} = \begin{bmatrix} p_{34}u - p_{14} \\ p_{34}v - p_{24} \\ p'_{34}u' - p'_{14} \\ p'_{34}v' - p'_{24} \end{bmatrix}$$

• Where p'_{11} etc. are the parameters inside the camera projection matrix for camera 2 (O'), and w' = (u', v') are the pixel coordinates of X projected on the image plane I' of the second camera











• Using an additional camera has forced **A** to become over-constrained.

$$\begin{bmatrix} p_{11} - p_{31}u & p_{12} - p_{32}u & p_{13} - p_{33}u \\ p_{21} - p_{31}v & p_{22} - p_{32}v & p_{23} - p_{33}v \\ p'_{11} - p'_{31}u' & p'_{12} - p'_{32}u' & p'_{13} - p'_{33}u' \\ p'_{21} - p'_{31}v' & p'_{22} - p'_{32}v' & p'_{23} - p'_{33}v' \end{bmatrix} \begin{bmatrix} X \\ Y \\ Z \end{bmatrix} = \begin{bmatrix} p_{34}u - p_{14} \\ p_{34}v - p_{24} \\ p'_{34}u' - p'_{14} \\ p'_{34}v' - p'_{24} \end{bmatrix}$$

• Therefore we can now find a least-squares solution:

$$Ax = b$$
$$\Leftrightarrow x = (A^{T}A)^{-1}A^{T}b$$







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Beyond triangulation

We have seen the simplest form of stereo vision: Given a pair of *calibrated* cameras observing a single feature at *corresponding* pixel locations $w \leftrightarrow w'$, the 3D position of the corresponding world location X can be estimated via triangulation











Beyond triangulation

How is the correspondence problem solved if there are several points $\{w_i\}_{i=1}^{N_1}$ in image 1 and several points $\{w'_j\}_{j=1}^{N_2}$ in image 2?











Beyond triangulation

SIFT will give us a set of proposed correspondences $\{w_i \leftrightarrow w'_j\}$ but there will be **many** outliers in these proposals

The question is then how can we remove outliers ?

Using the **epipolar constraint**









Epipolar Geometry

- To understand the **epipolar constraint**, we first need to understand the **geometry** that relates the
 - cameras
 - points in 3D space
 - and their corresponding observations $\{w_i \leftrightarrow w'_j\}$



• This type of geometry is referred to as the **epipolar geometry**





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Epipolar Geometry - fundamentals

Lets revisit our stereo pair











The baseline l_b is the line joining the two optical centers.











The epipolar plane Π is the plane defined by the 3D point X and the optical centers of the cameras.











An **epipole** is the point of intersection of the **baseline** with the image plane. There are two epipoles e and e', one for each image.











An epipolar line is a line of intersection of the epipolar plane with an image plane. It is the image, in one camera, of the ray from the other camera's optical centre to the point X.











For different world points X, the epipolar plane rotates about the baseline. All epipolar lines intersect at their corresponding epipole.











The **epipolar constraint** limits the search for correspondences, from the region of the whole image, to only the pixels spanned by the epipolar line.











If a point feature w is observed in one image, then its location w' in the other image must lie on its corresponding epipolar line l'











So a simple algorithm for determining correspondences is to match each feature w_i from camera 1 to a feature w'_j from camera 2 which is close to the epipolar line of camera 2, provided that the SIFT descriptors for w_i and w'_j are similar.











Therefore, we must calculate the equation of the epipolar lines.





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• Firstly, lets assume that the 1^{st} camera is located at the world origin $oldsymbol{O}$

• This means that every point X is expressed in the camera coordinates of the 1st camera, since its optical center coincides with the world origin $\Rightarrow X \rightarrow X_c$









• If this is the case, the 2nd camera is at a location **O'** in world coordinates, which can be expressed by a **translation t** and **rotation R**, w.r.t. 1st camera, i.e., the world origin











- Every 3D point X can be expressed as X_c , which is the cameracentered coordinate system of the 1st camera, and as X'_c , which is the camera-centered coordinate system of the 2nd camera.
- Since $X = X_c$, we can relate $X \leftrightarrow X'_c$ or $X_c \leftrightarrow X'_c$ using a Euclidean transformation composed by the translation t and rotation R of the 2nd camera, w.r.t. the 1st camera









Here is how to find an expression for the epipolar line:



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This can be rewritten in matrix form:

 $X'_c \cdot (t \times RX_c) = \mathbf{0}$ $\Leftrightarrow X'^T_c EX_c = \mathbf{0}$

Where $E = T_{\times}R$ is the **essential matrix,** and

$$\boldsymbol{T}_{\times} = \begin{bmatrix} 0 & -t_z & t_y \\ t_z & 0 & -t_x \\ -t_y & t_x & 0 \end{bmatrix}$$

is a matrix representing the cross product with t such that $\mathbf{t} imes \boldsymbol{\nu} = T_{ imes} \mathbf{v}$







Fundamental Matrix

Recall that for $X_c \leftrightarrow w$: $\widetilde{w} = KX_c \Leftrightarrow X_c = K^{-1}\widetilde{w}$

and similarly, for $X'_c \leftrightarrow w'$: $\widetilde{w}' = K'X'_c \Leftrightarrow X'_c = K'^{-1}\widetilde{w}'$

Combining the two equations yields the equation of the two epipolar lines in pixel coordinates: $\mathbf{v}'^T \mathbf{F} \mathbf{v} = 0$

$$X_{c}^{\prime T} \mathbf{E} \mathbf{X}_{c} = 0$$

$$\Rightarrow (\mathbf{K}^{\prime - 1} \widetilde{\mathbf{W}}^{\prime})^{T} \mathbf{E} (\mathbf{K}^{-1} \widetilde{\mathbf{W}}) = 0$$

$$\Rightarrow \widetilde{\mathbf{W}}^{\prime T} (\mathbf{K}^{\prime - T} \mathbf{E} \mathbf{K}^{-1}) \widetilde{\mathbf{W}} = 0$$

$$\Rightarrow \widetilde{\mathbf{W}}^{\prime T} \mathbf{F} \widetilde{\mathbf{W}} = \mathbf{0}$$

where $F = K'^{-T} E K^{-1}$ is the **fundamental matrix**.









Fundamental Matrix

This is the equation of the epipolar line in either camera:

 $\widetilde{\boldsymbol{w}}^{\prime T}\boldsymbol{F}\,\widetilde{\boldsymbol{w}}=\boldsymbol{0}$

Assuming we know the fundamental matrix, for every point \boldsymbol{w} in image 1, this expression gives us the line in image 2 on which the corresponding \boldsymbol{w}' must lie, and vice versa.

- $l' = F\widetilde{w}$ is the epipolar line in the 2nd image, associated with w
- $l = F^T \widetilde{w}'$ is the epipolar line in the 1st image, associated with w'
- l_i : ax + by + c = 0









Here are a few examples of the epipolar constraint













Epipolar constraint examples: Parallel image planes



- Baseline intersects the image planes at infinity
- Epipoles are at infinity
- Epipolar lines are parallel to the *u*-axis of each image plane









Epipolar constraint examples: Parallel image planes









Epipolar constraint examples: Forward translation



• The **epipoles** have the same position in both images





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Examples

Epipolar constraint examples: Forward translation













Fundamental Matrix

In order to apply the epipolar constraint we need to know the fundamental matrix:

 $F = K'^{-T} E K^{-1}$ $= K'^{-T} T_{\times} R K^{-1}$

All the parameters of F come from the calibration of the two cameras.

Remember that the perspective camera model $P_{ps} = K[R|T]$ contains this information.

If, however, these are not available (e.g., because we used a projective camera model to calibrate our cameras), **F** must be estimated using known image correspondences.











To estimate the fundamental matrix, we follow a similar approach as in camera calibration

$$\widetilde{\boldsymbol{w}}'^{T}\boldsymbol{F}\,\widetilde{\boldsymbol{w}} = 0 \Rightarrow \begin{bmatrix} u' \, v' \, 1 \end{bmatrix} \begin{bmatrix} f_{11} & f_{12} & f_{13} \\ f_{21} & f_{22} & f_{23} \\ f_{31} & f_{32} & f_{33} \end{bmatrix} \begin{bmatrix} u \\ v \\ 1 \end{bmatrix} = 0$$

Each point correspondence $w \leftrightarrow w'$ generates a single equation for estimating the parameters inside F.









Here are N such correspondences:

$$\begin{bmatrix} u_{1}u'_{1} & v_{1}u'_{1} & u'_{1} & u_{1}v'_{1} & v_{1}v'_{1} & v'_{1} & u_{1} & v_{1} \\ & & \vdots & & & \\ u_{N}u'_{N} & v_{N}u'_{N} & u'_{N} & u_{N}v'_{N} & v_{N}v'_{N} & v'_{N} & u_{N} & v_{N} \end{bmatrix} \begin{bmatrix} f_{11} \\ f_{12} \\ f_{13} \\ f_{21} \\ f_{22} \\ f_{23} \\ f_{31} \\ f_{32} \end{bmatrix} = \begin{bmatrix} -1 \\ \vdots \\ -1 \end{bmatrix}$$

The minimal number for N is 8, because **F** has 8 degrees of freedom (we can set $f_{33} = 1$)







• In practise, 8 clearly indicated correspondences are never available

• What is available is a set of correspondences proposed by SIFT, for a large set of features, in which there are **always** errors









- We can use RANSAC to solve this problem
 - 1. Obtain 8 random SIFT correspondences. Use them to estimate **F**.
 - 2. For every feature w_i in image 1 calculate the epipolar line in image 2. Check if the corresponding feature w'_j in image 2 proposed by SIFT falls "close" to the epipolar line. Count the number *S* of such "inliers".
 - 3. If $S \ge T$ where T is a threshold, then there is consensus with the random sample taken in the first step. Calculate **F** for all inliers and terminate here.
 - 4. If S < T then no consensus is reached. Repeat from step 1.
 - 5. If after N iterations no consensus is reached, select the model that gave the highest S, calculate F using all inliers in S and terminate







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Thank you.



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